



IllustrisTNG

What's Next for the EFTofLSS*?

* Effective Field Theory of Large Scale Structure

Oliver Philcox (Princeton)

UK Cosmology Meeting, September 2020

Based on: [2002.04035](#), [2008.08084](#), [2004.09515](#) & [2006.10055](#)



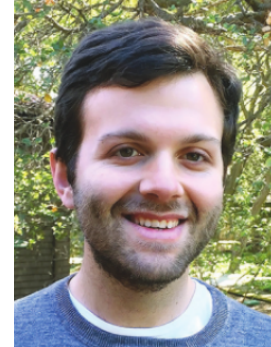
Elena Massara



Francisco Villaescusa-Navarro



David Spergel



Blake Sherwin



Gerrit Farren



Eric Baxter



Mikhail Ivanov



Marcel Schmittfull



Marko Simonovic



Matias Zaldarriaga

EFT of Matter: A Lightning Introduction

Standard Perturbation Theory [e.g. Bernardeau+02]

- Solve the **ideal** fluid equations
- Expansion variable: overdensity field $\delta(\mathbf{x})$
- Assume valid on **all** scales

$$\begin{aligned}\dot{\delta}_\Lambda + \nabla \cdot [(1 + \delta_\Lambda) \mathbf{v}_\Lambda] &= 0 \\ \dot{\mathbf{v}}_\Lambda + (\mathbf{v}_\Lambda \cdot \nabla) \mathbf{v}_\Lambda &= -\mathcal{H} \mathbf{v}_\Lambda - \nabla \phi_\Lambda - \frac{1}{\rho_\Lambda} \nabla \underline{\underline{\tau}}\end{aligned}$$

Effective Field Theory [e.g. Carrasco+12, Baumann+12]

- Solve the **non-ideal** fluid equations, including **viscosity** etc.
- Expansion variable: **smoothed** overdensity field $\delta_\Lambda(\mathbf{x})$
- Theory **only** valid for $k < \Lambda$
- EFT includes:
 - Backreaction of small-scale physics on large-scale modes
 - **Controlled** ultraviolet (UV) behavior
 - Proper treatment of **non-perturbative** long-wavelength displacements [e.g. Senatore+14]

$$P(k) = \underbrace{P_{\text{lin}}(k) + P_{22,\Lambda}(k) + 2P_{13,\Lambda}(k)}_{\text{1-loop SPT}} - \underbrace{2c_{s,\Lambda}^2 k^2 P_{\text{lin}}(k)}_{\text{1-loop EFT}}$$

At one-loop, EFT adds a **free parameter** c_s^2 encoding **small-scale physics** that must be **fit** to data

EFT of Matter: A Lightning Introduction

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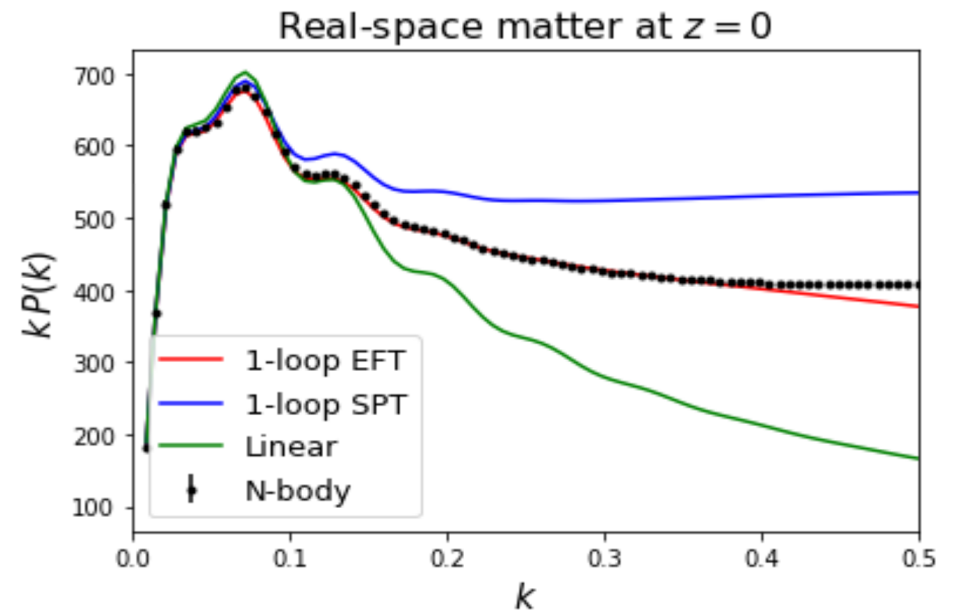
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Biased Tracers

- **Matter** is not observed directly → must predict **galaxy** statistics
- Simple approach: expand the galaxy overdensity in powers of δ :

$$\delta_g(\mathbf{x}) = b_1 \delta(\mathbf{x}) + \frac{b_2}{2} \delta^2(\mathbf{x}) + \frac{b_3}{6} \delta^3(\mathbf{x}) + \dots$$

- The EFT approach: include all possible parameters allowed by symmetry

$$\delta_g = b_1 \delta + \epsilon + \frac{b_2}{2} \delta^2 + b_{\mathcal{G}_2} \mathcal{G}_2 + \frac{b_3}{6} \delta^3 + b_{\delta \mathcal{G}_2} \delta \mathcal{G}_2 + b_{\mathcal{G}_3} \mathcal{G}_3 + b_{\Gamma_3} \Gamma_3 + R_*^2 \partial^2 \delta$$

with **density operators**, **tidal operators**, **stochastic operators**, and **non-local operators** (all integrated over a lightcone)

- Also needs **redshift-space** distortions: $\mathbf{s} = \mathbf{r} + \frac{\mathbf{v} \cdot \hat{\mathbf{z}}}{aH} \hat{\mathbf{z}}$

Biased Tracers

- Full 1-loop **EFT** model for the **redshift-space** power spectrum of **biased tracers** [e.g. Perko+16]

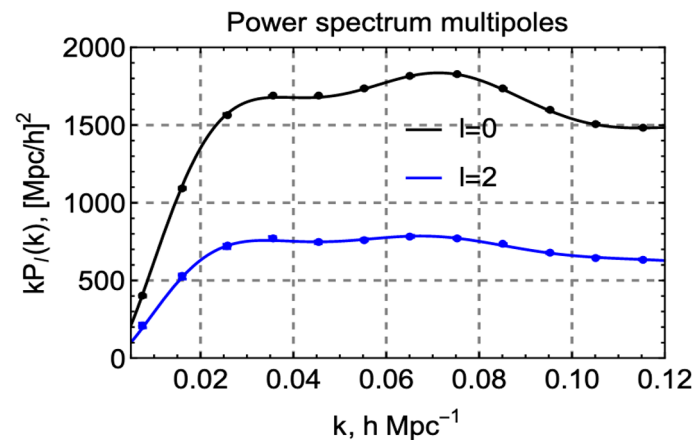
$$P_{g,\ell}(k) = P_{g,\ell}^{\text{tree}}(k) + P_{g,\ell}^{1\text{-loop}}(k) + P_{g,\ell}^{\text{noise}}(k) + P_{g,\ell}^{\text{ctr}}(k)$$

Linear Theory 1-loop SPT Stochastic Terms Counterterms

- **Seven** nuisance parameters:

$$\{b_1, b_2, b_{G_2}, c_{S,0}, c_{S,2}, b_4, P_{\text{shot}}\} \times \{\omega_b, \omega_{\text{cdm}}, h, A_S, \sum m_\nu, n_s, \dots\}$$

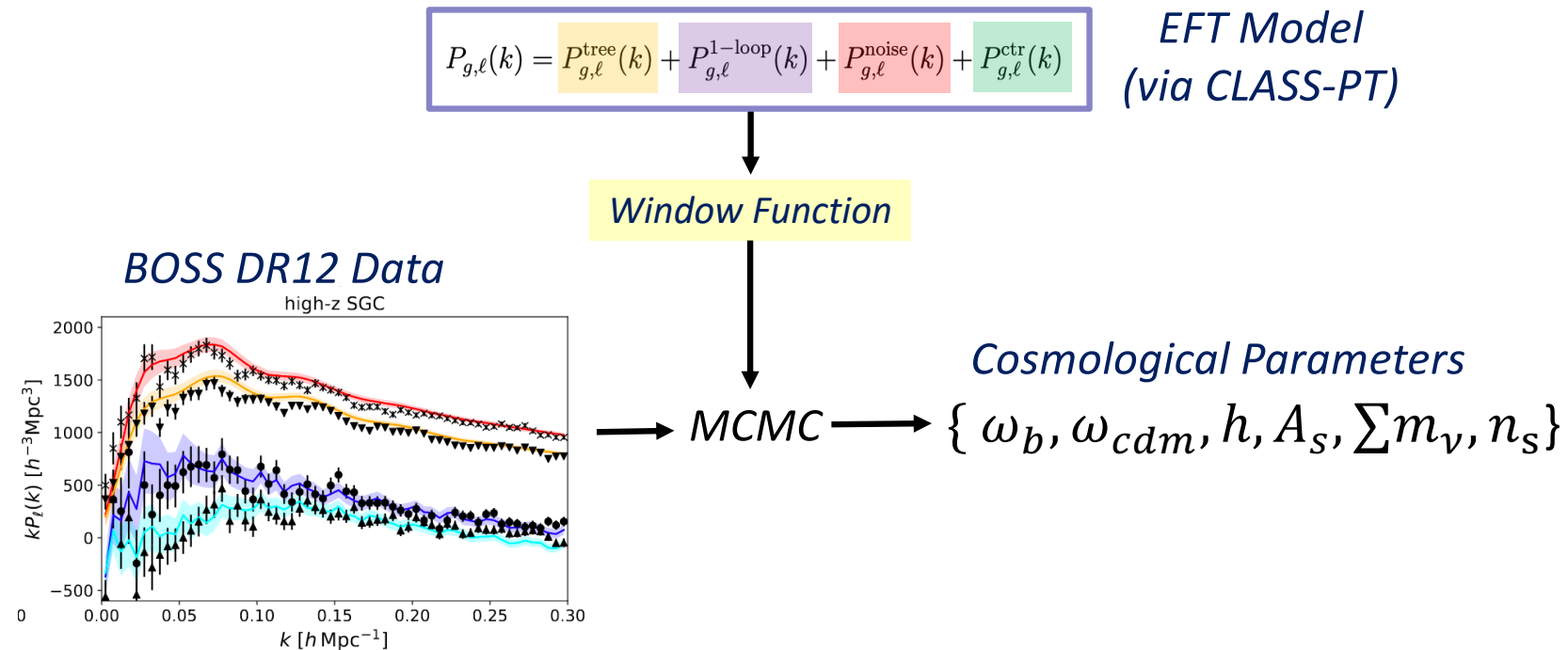
- **Sub-percent** accurate on large-scales



Beyond the Blackboard

<https://github.com/michalychforever/CLASS-PT>

- Can we use EFT to fit the **full-shape** of **observed** galaxy power spectra?



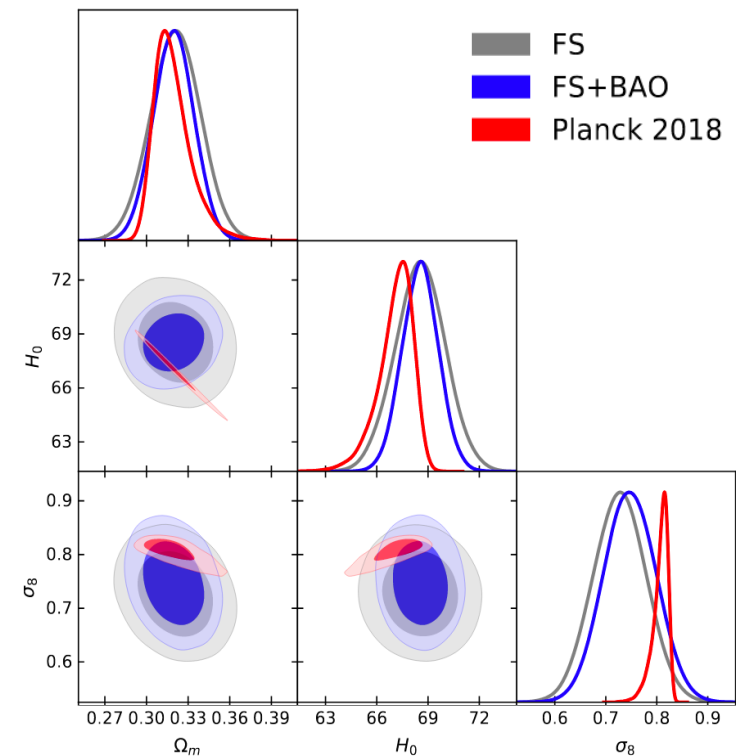
Ivanov+19a,b, d'Amico+19, Chudaykin+20, Philcox+20a

Cosmology from Galaxies + EFT

- Get **competitive** constraints on cosmology from **galaxy** power spectra plus **BBN** [Ivanov+19]
- Constraints are **strengthened** by combining with extra **BAO** information from *reconstruction* [Philcox+20a]

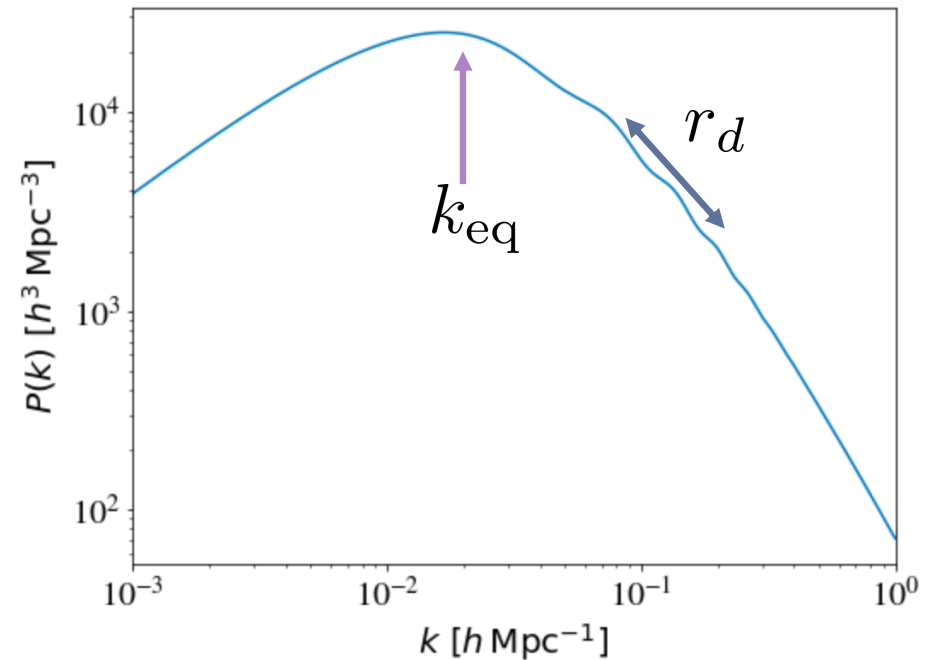
H_0	BOSS FS+BBN	BOSS FS+BAO+BBN	Planck 2018
	68.55 ± 1.5	67.9 ± 1.1	$67.1^{+1.3}_{-0.72}$

- As strong H_0 constraints as *Planck* **without** using CMB data!



H_0 without the Sound Horizon

- H_0 information comes from two **standard rulers**:
 1. Matter-radiation **equality** ($z \sim 3600$)
 2. **Baryon Acoustic Oscillation** scale ($z \sim 1100$)
- Can **isolate** the **equality** information by **removing** the BBN prior on ω_b in the EFT analysis

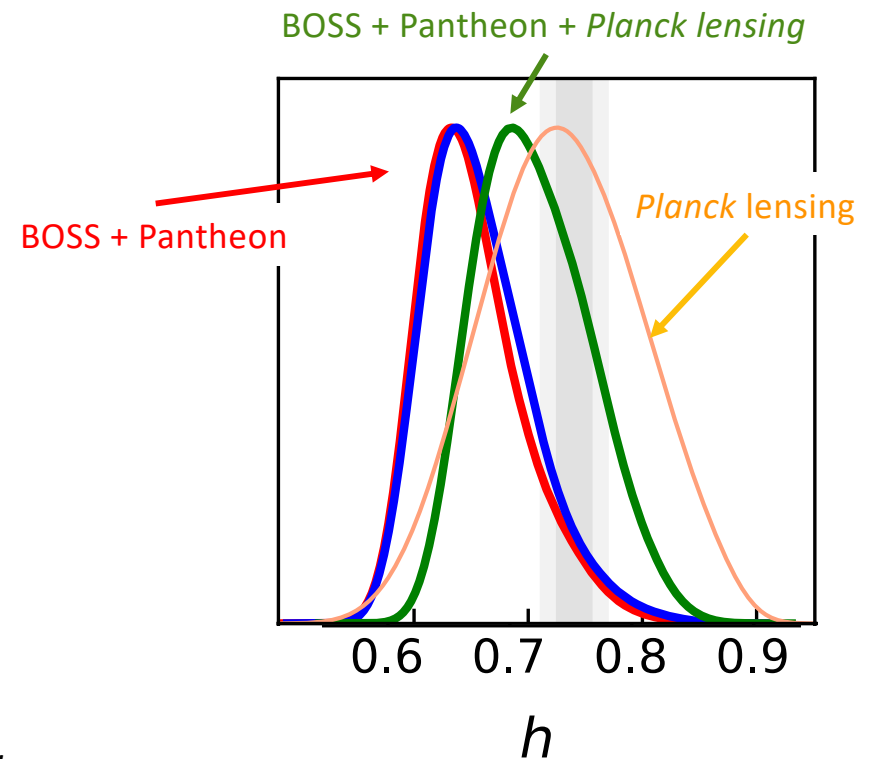


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H_0	BOSS + Pantheon	+ Planck lensing
	$65.1^{+3.0}_{-5.4}$	$70.6^{+3.7}_{-5.1}$

- Galaxy constraints are **independent** of sound-horizon physics, and are $\sim 2\sigma$ below SHOES
- $\sigma_{H_0} \sim 1.5 \text{ km s}^{-1} \text{ Mpc}^{-1}$ will soon be possible with *Euclid*

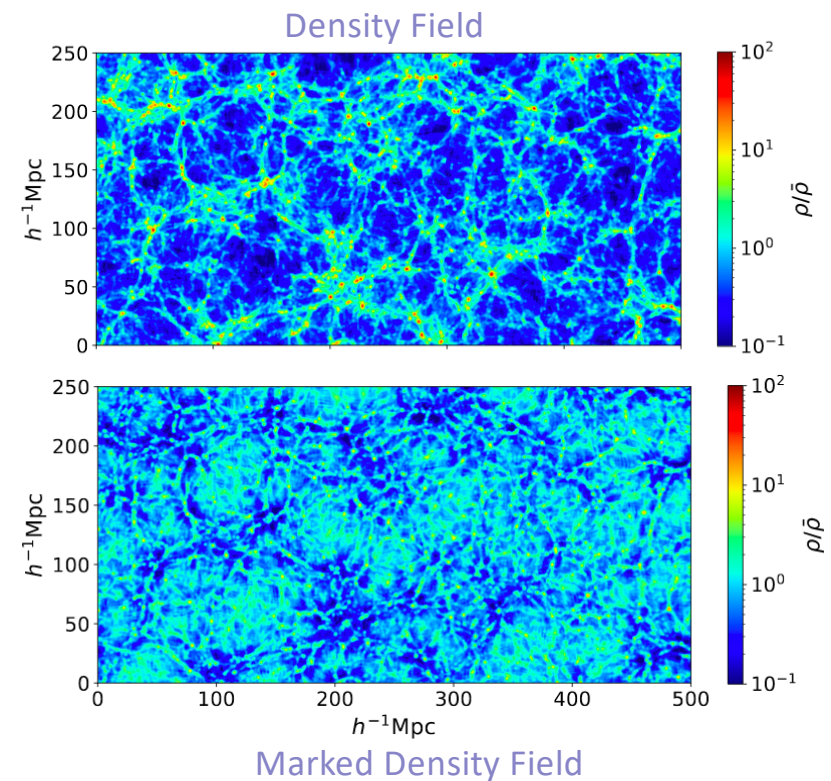


Alternative Density Statistics

- **Low-density regions** carry a lot of information but contribute little to δ [e.g. Pisani+19]
- Alternative statistics can help here, e.g. the **marked** density field [Stoyan 84, White 16, Massara+20]
- This is a **local-overdensity** weighted density field, shown to give strong constraints on the **neutrino** mass [Massara+20]

$$\rho_m(\mathbf{x}) \propto \frac{\rho(\mathbf{x})}{[c + \rho_R(\mathbf{x})]^p}$$

- Can we understand this using EFT?



White 16, Massara+20

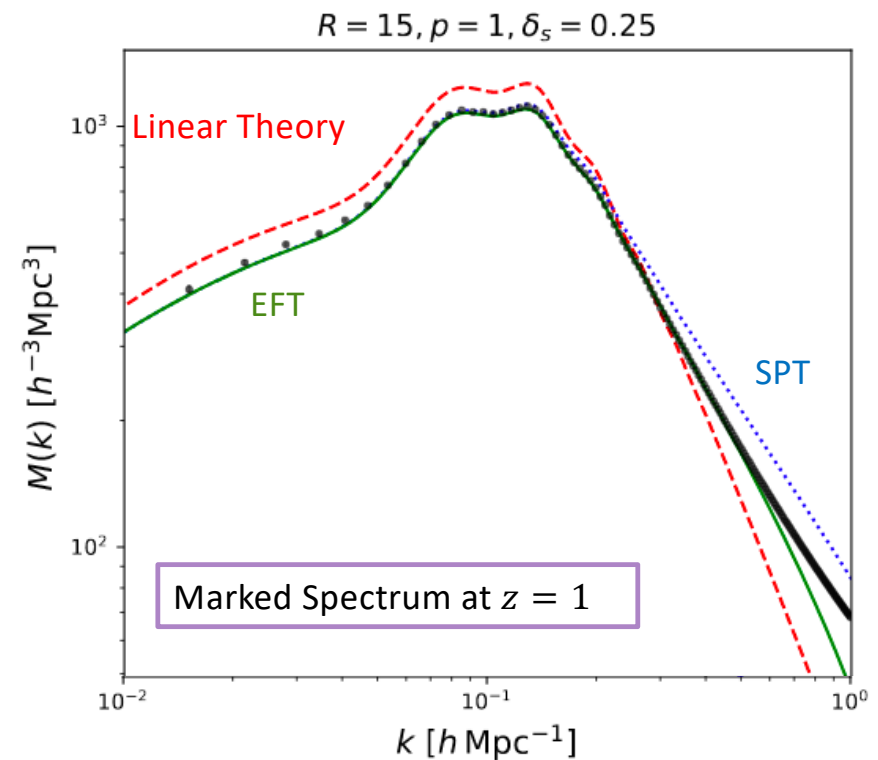
Understanding Marked Statistics

- Write down a **perturbative model** for the **marked** field

$$M(\mathbf{k}) = |\delta_M(\mathbf{k})|^2 = \frac{1}{\bar{m}^2} [M_{11}(\mathbf{k}) + M_{22}(\mathbf{k}) + 2M_{13}(\mathbf{k}) + 2M_{ct}(\mathbf{k})]$$

Linear Theory
1-loop SPT
Counterterms

- This theory is unusual:
 - **Linear** theory fails on **all** scales
 - At low redshifts, **EFT** fails on **all** scales
- **Small-scales** are coupled to **large scales**, through non-linearities and gravitational non-Gaussianities.
- This **shifts** small-scale information, e.g. about neutrinos and n_s , up to **quasi-linear scales**



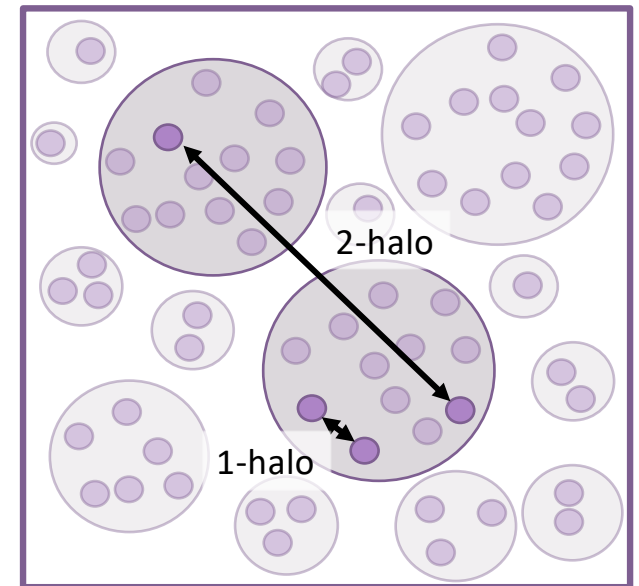
Philcox+20c, Philcox+ (in prep)

EFT + The Halo Model

- **EFT** gives a highly **accurate** $P(k)$ model at low k but **fails** close to the non-linear scale
- The **halo model** gives a **rough** model of $P(k)$ up to high k
- In **combination**, we can get **accurate** models up to **high- k**

$$P(k) = \overbrace{[I_1^1(k)]^2}_{\text{2-halo}} \overbrace{P_L(k)}_{\text{1-halo}} + \overbrace{I_2^0(k)}_{\text{1-halo}}$$

Mass Integrals Linear Power



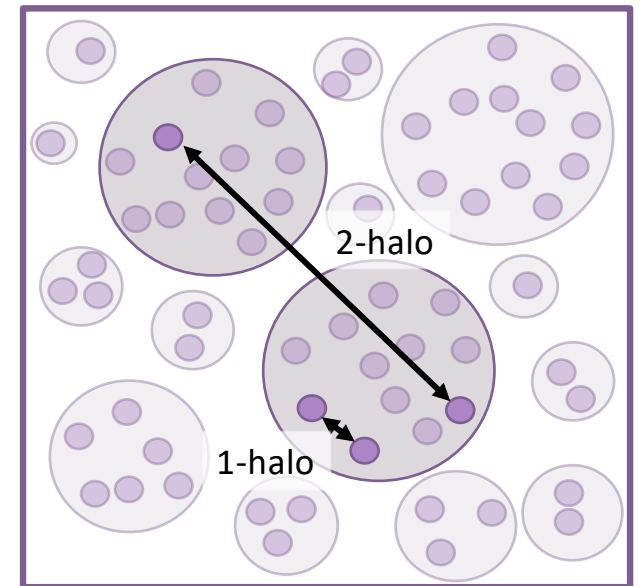
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$$P(k) = \underbrace{[I_1^1(k)]^2}_{\text{Mass Integrals}} \underbrace{P_{\text{EFT}}(k)}_{\text{EFT Power}} \underbrace{W^2(kR)}_{\text{Smoothing Function}} + \underbrace{I_2^0(k)}_{\text{1-halo}}$$

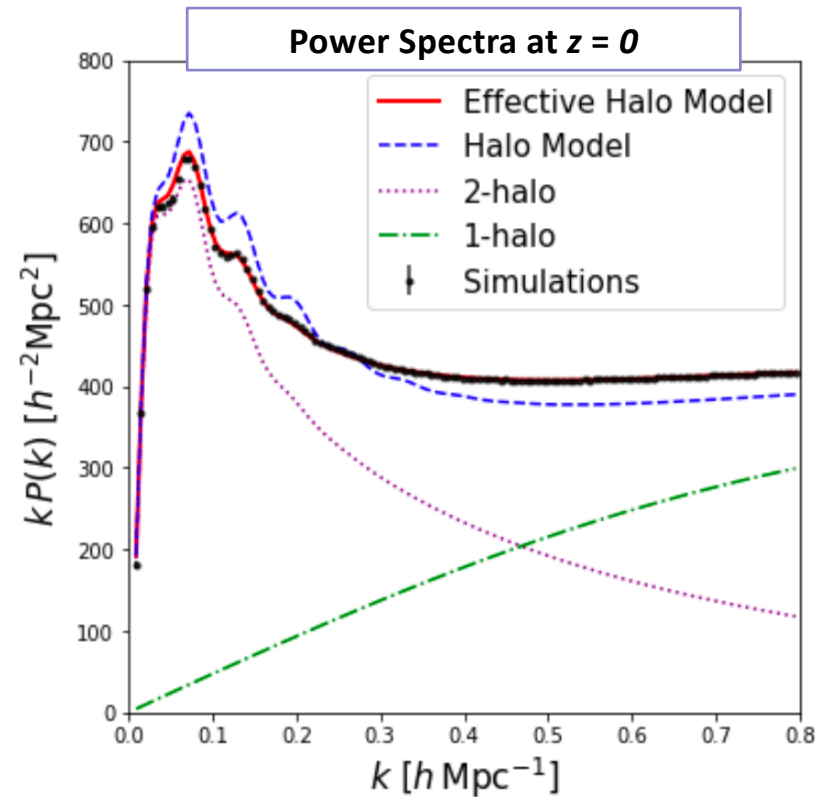
2-halo

1-halo



The Effective Halo Model

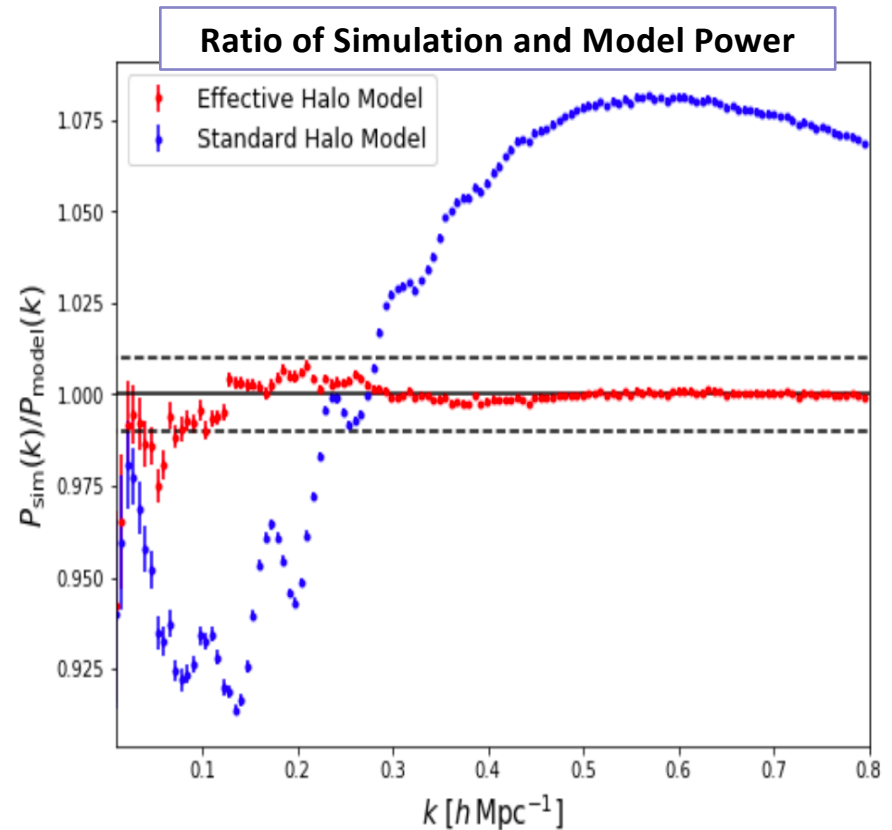
- The Effective Halo Model is based on **perturbation theory** and is **1% accurate** for a large range of cosmologies.



Philcox+ 2020b, Philcox+ (in prep.)

The Effective Halo Model

- The Effective Halo Model is based on **perturbation theory** and is **1% accurate** for a large range of cosmologies.
- We can also predict **covariances** between halo **number counts** and $P(k)$.
- This can be used to compute projected spectra e.g.
 - **Weak Lensing**
 - Joint analysis of lensing and **thermal SZ**



Philcox+ 2020b, Philcox+ (in prep.)

A visualization of the cosmic web, showing a complex network of blue filaments and nodes with numerous bright orange and yellow galaxies scattered throughout. The background is dark blue/black.

arXiv:

[2002.04035](#)

[2008.08084](#)

[2004.09515](#)

[2006.10055](#)

More questions?

Email ohp2@cantab.ac.uk

Summary

The *EFTofLSS*:

- Gives **accurate** models for density statistics
- Allows for **parameter inference** shedding light on H_0 tension
- Helps us to **understand** alternative statistics
- Can be combined with the **halo model** to accurately predict **all scales**

Many more possibilities, e.g.:

- Bispectra
- Field-level EFT
- Other statistics
- Weak lensing